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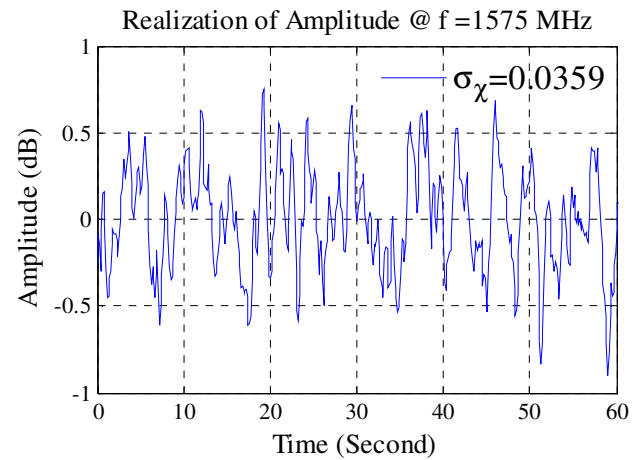
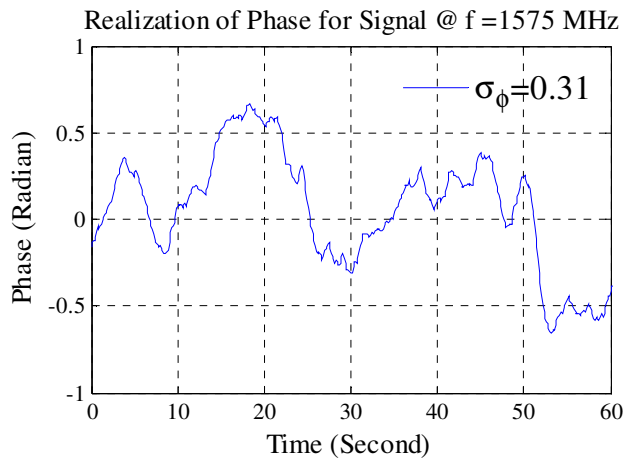
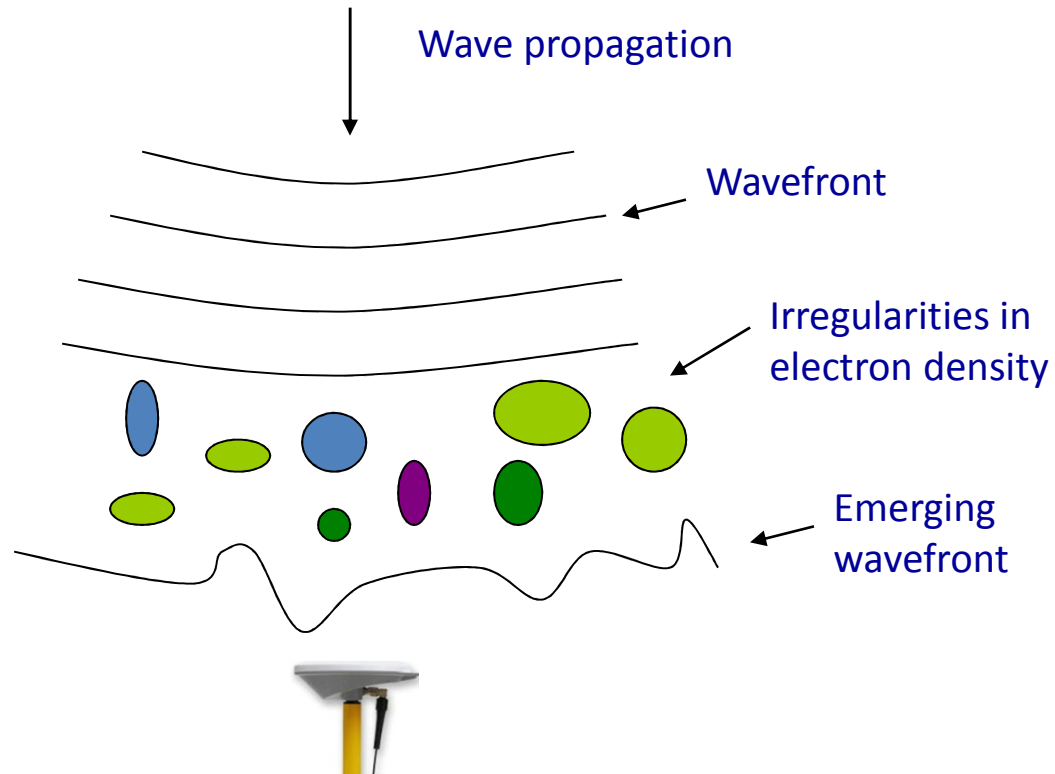
High Latitude Scintillation Analysis for Marine and Aviation Applications

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University of Calgary

Outline

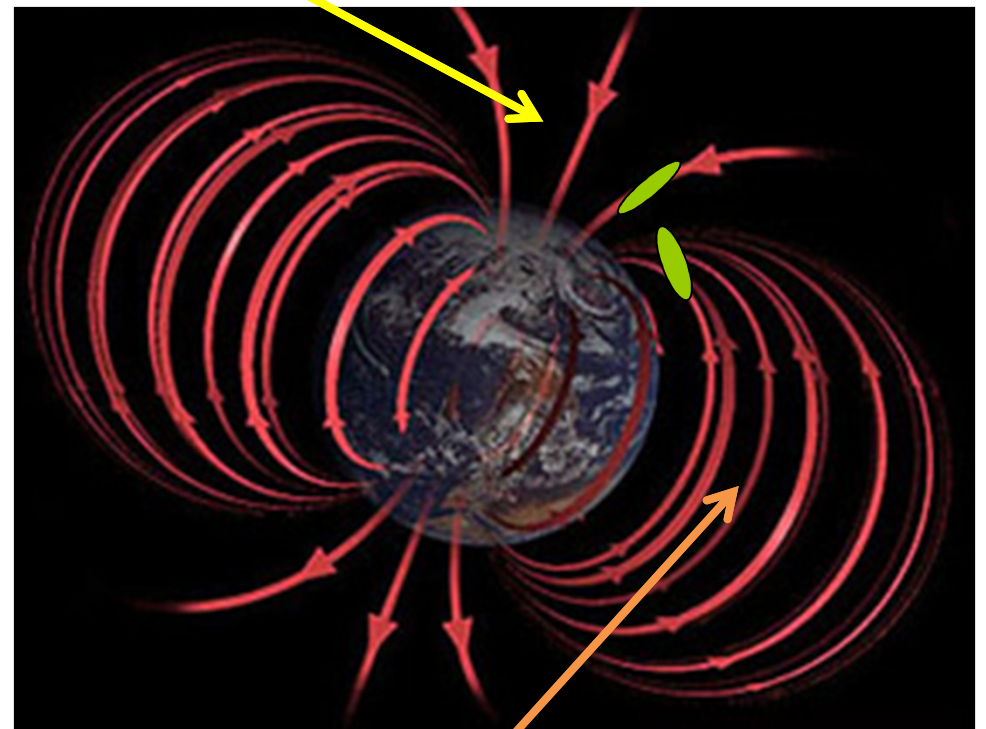
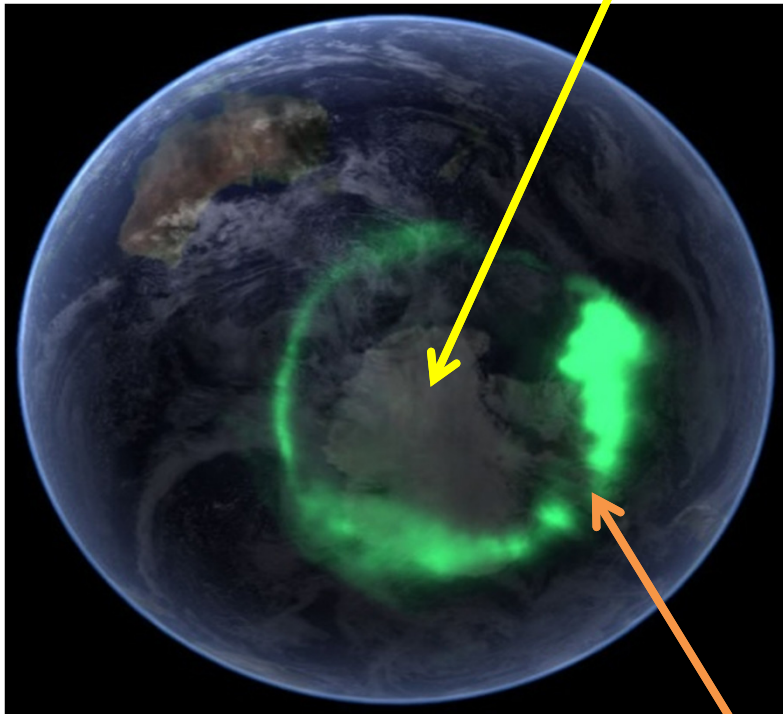
- Auroral/polar scintillation mechanisms
- Scintillation data set
- Simulation tools
- Results
- Summary

Ionospheric Scintillation



High Latitude Scintillations

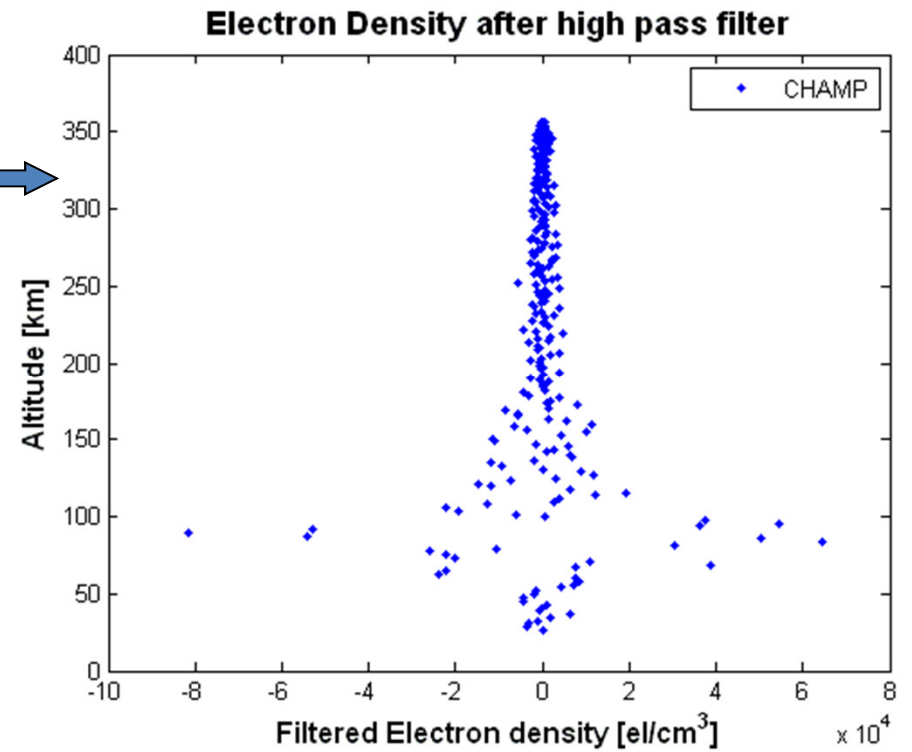
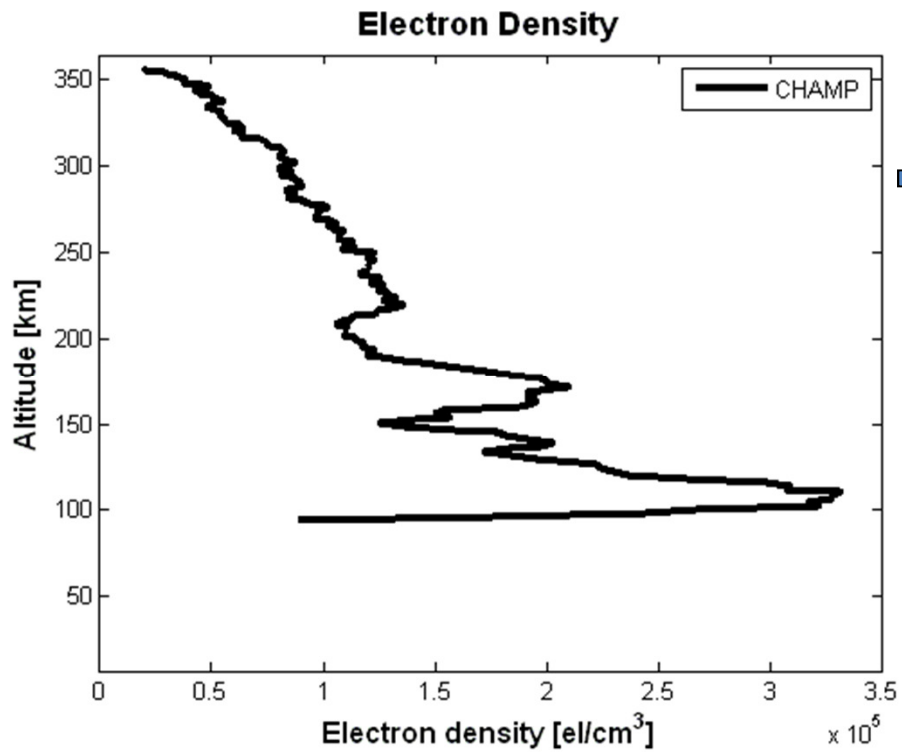
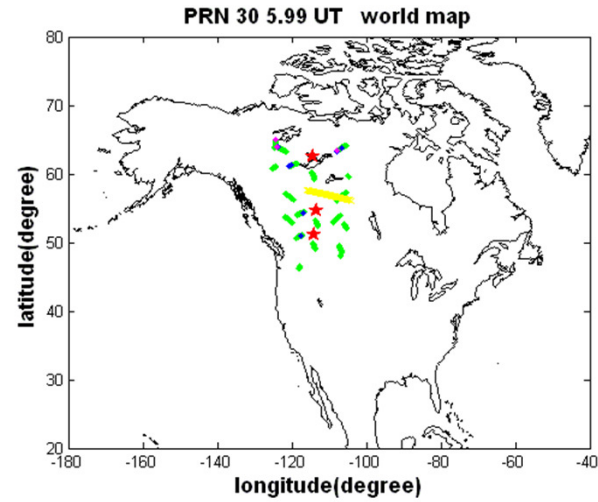
Polar Cap



Auroral Oval

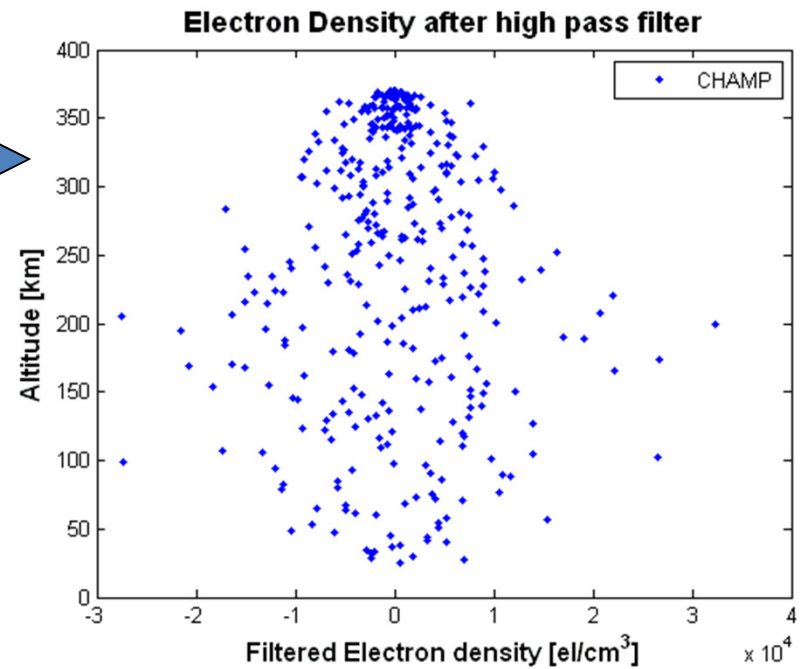
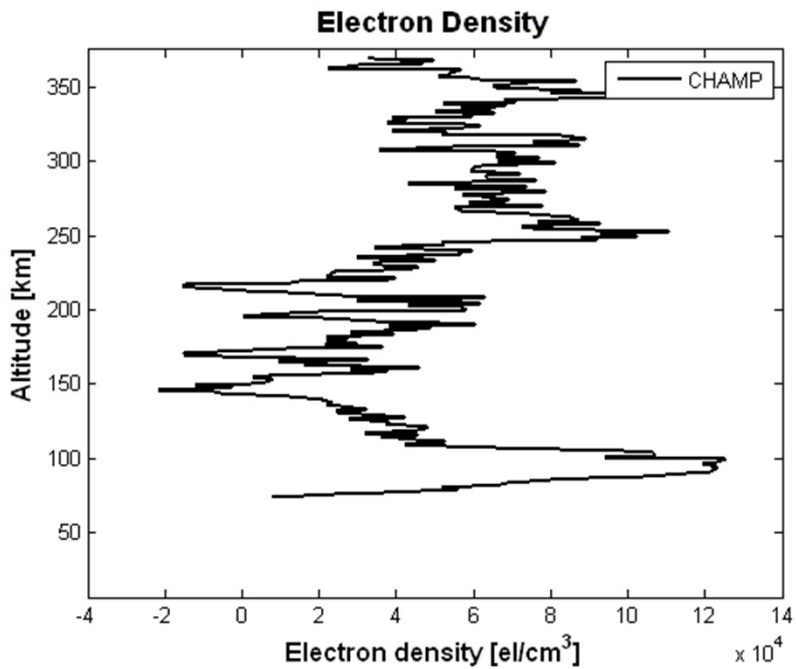
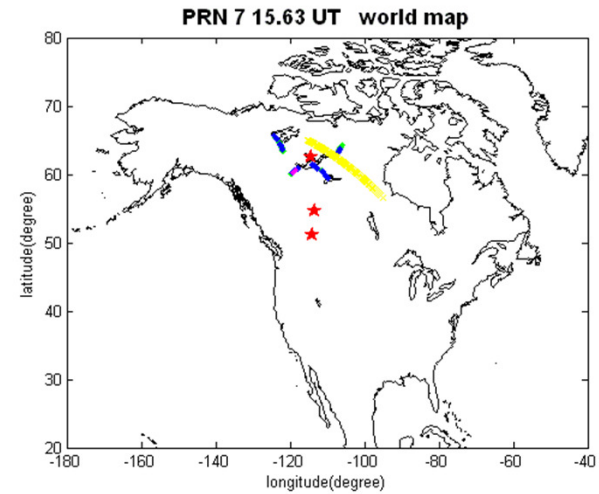
Auroral Irregularities

E-region:
October 1, 2006



Polar Irregularities

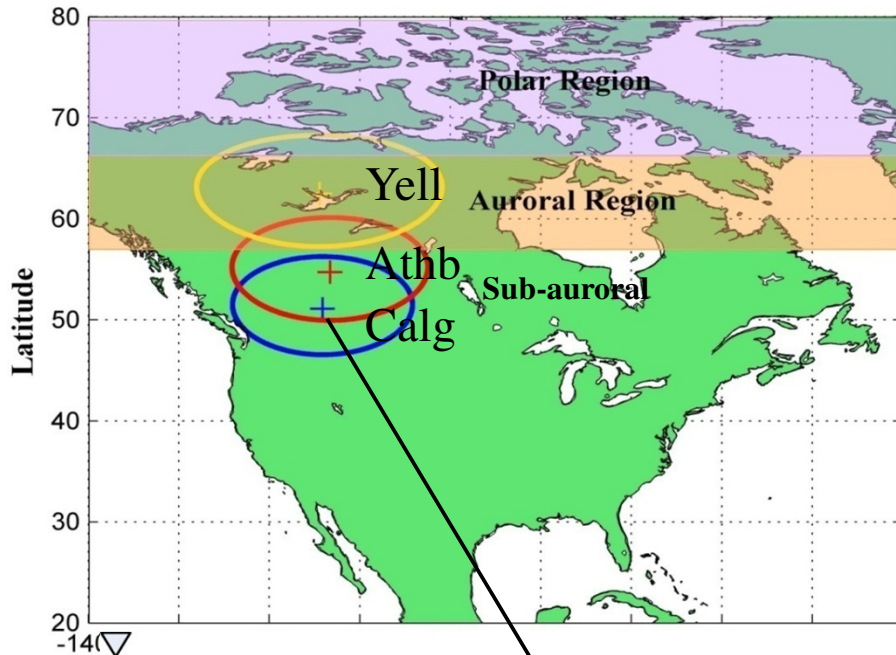
E- & F-region:
January 7, 2005



CANGIM Database

The CANGIM receivers measure

- 50 Hz raw GPS observations
- Phase scintillation index (σ)
- Amplitude scintillation index (S4)



Station	Geographic		Geomagnetic	
	Lat.	Long.	Lat.	Long.
Yell	62.48°	-114.48°	68.80°	298.37°
Athb	54.72°	-113.31°	61.52°	305.67°
Calg	51.08°	-114.13°	57.86°	306.43°



Note: - The coverage areas for IPP above 20° are indicated as circles.

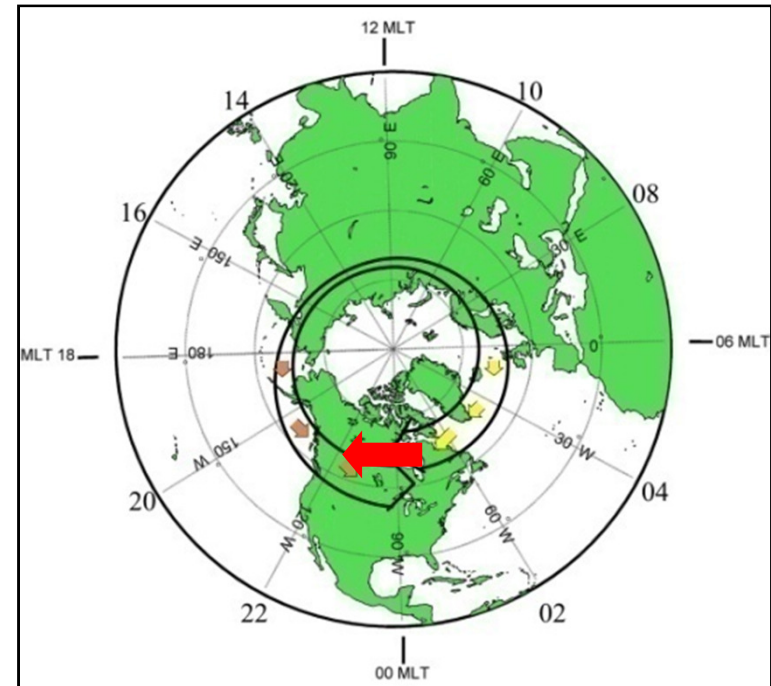
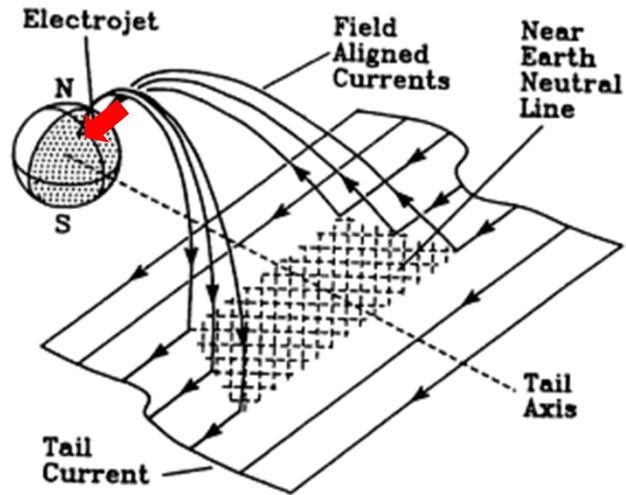
Objective

To develop flexible tools to study high latitude scintillation impact on receiver performance for various user groups (marine, aviation)

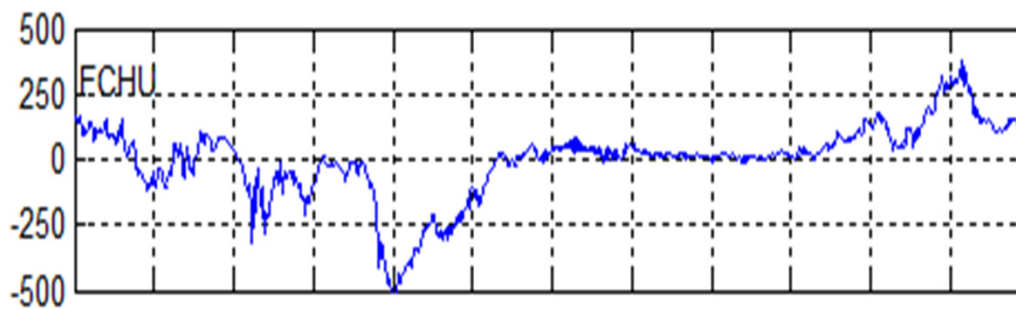
Method

- Determine scintillation characteristics for auroral and polar regions (CANGIM)
- Develop simulation based on observed high latitude characteristics
- Develop flexible tracking loops to assess impact – determine thresholds for loss of lock
- Quantify duration and frequency of receiver impact, number of satellites affected, etc.

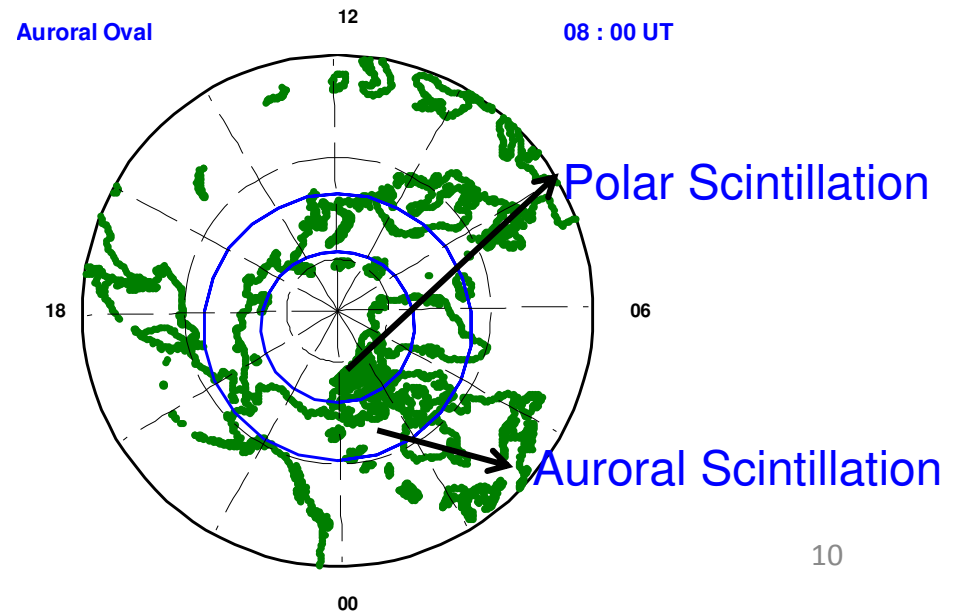
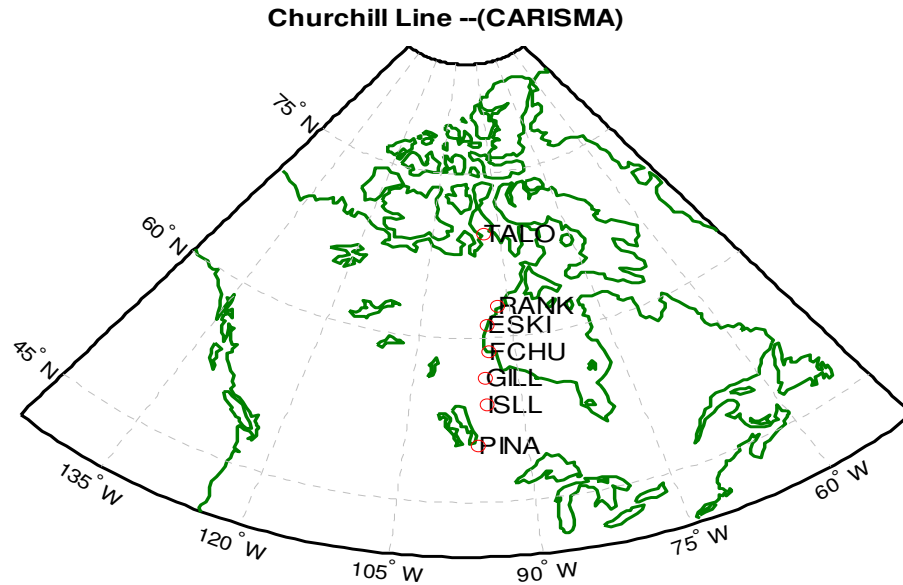
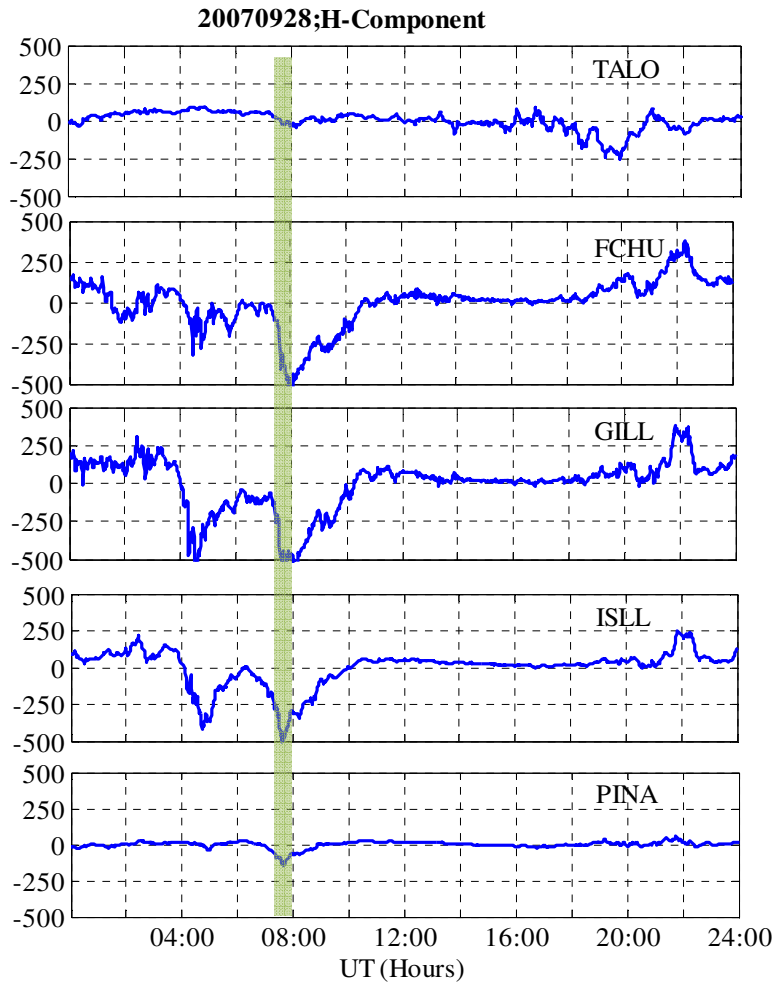
Auroral Substorm Current Wedge



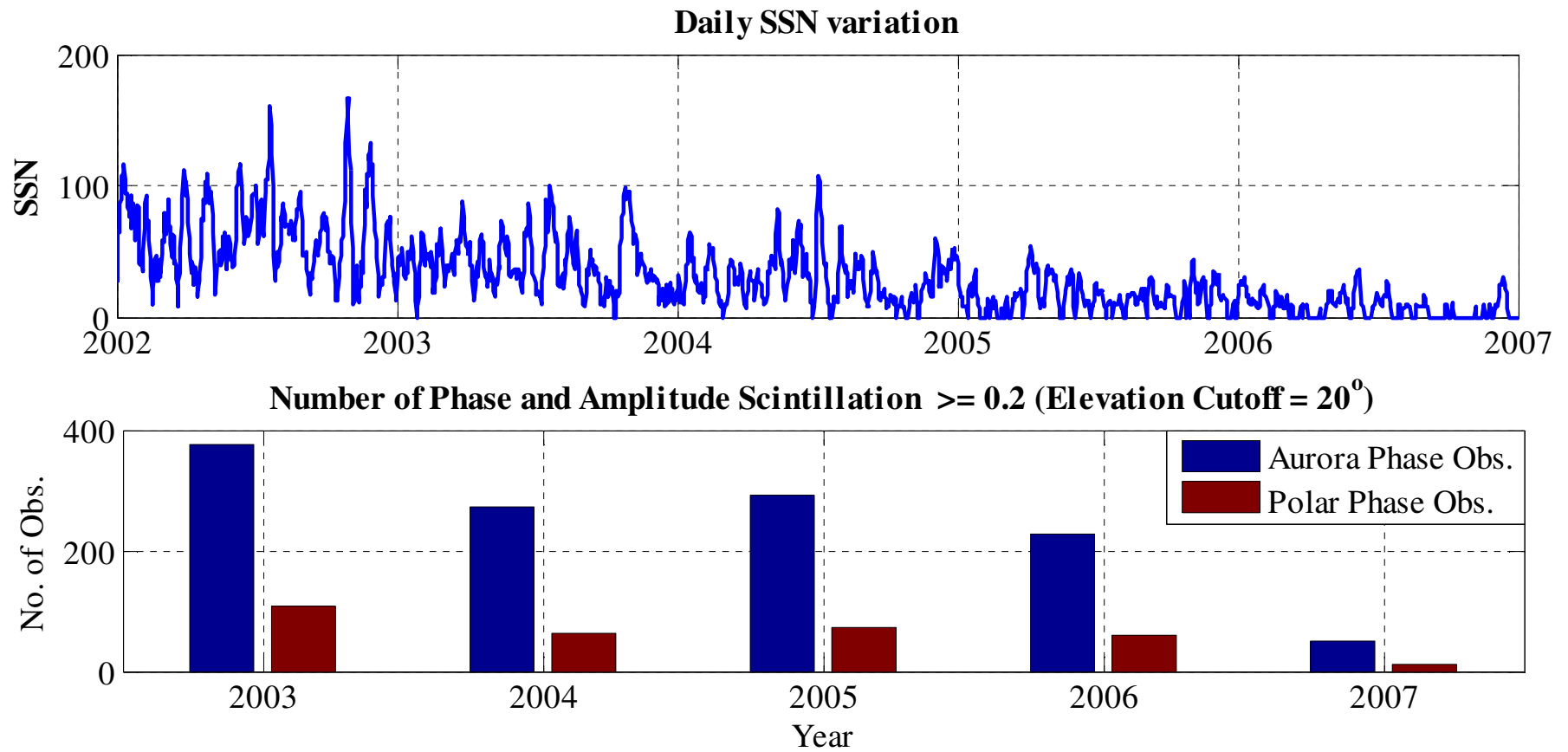
Magnetometer signature



Auroral Boundary Estimation

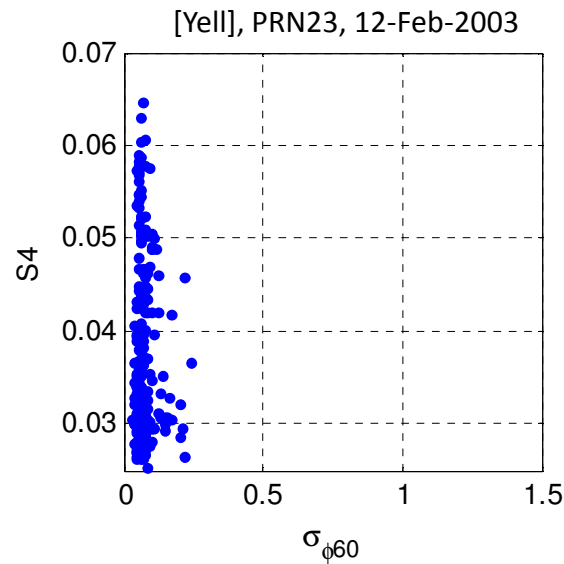


Event Selection

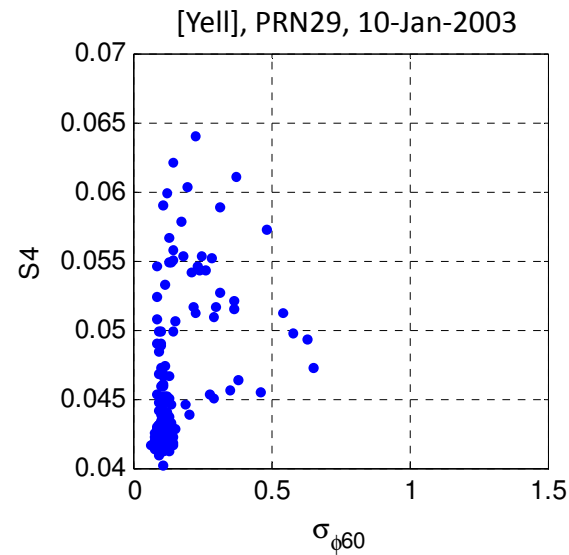


Auroral Scintillation Characteristics

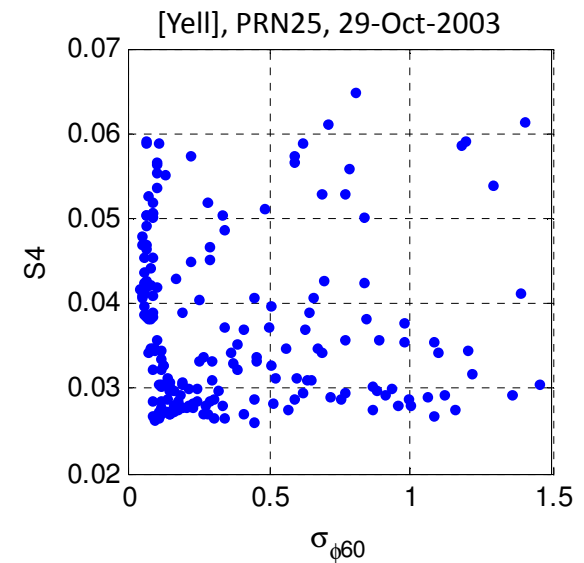
- Phase scintillation is dominant
- Low amplitude scintillation ($S4 < 0.1$) for all events



Weak Scintillation



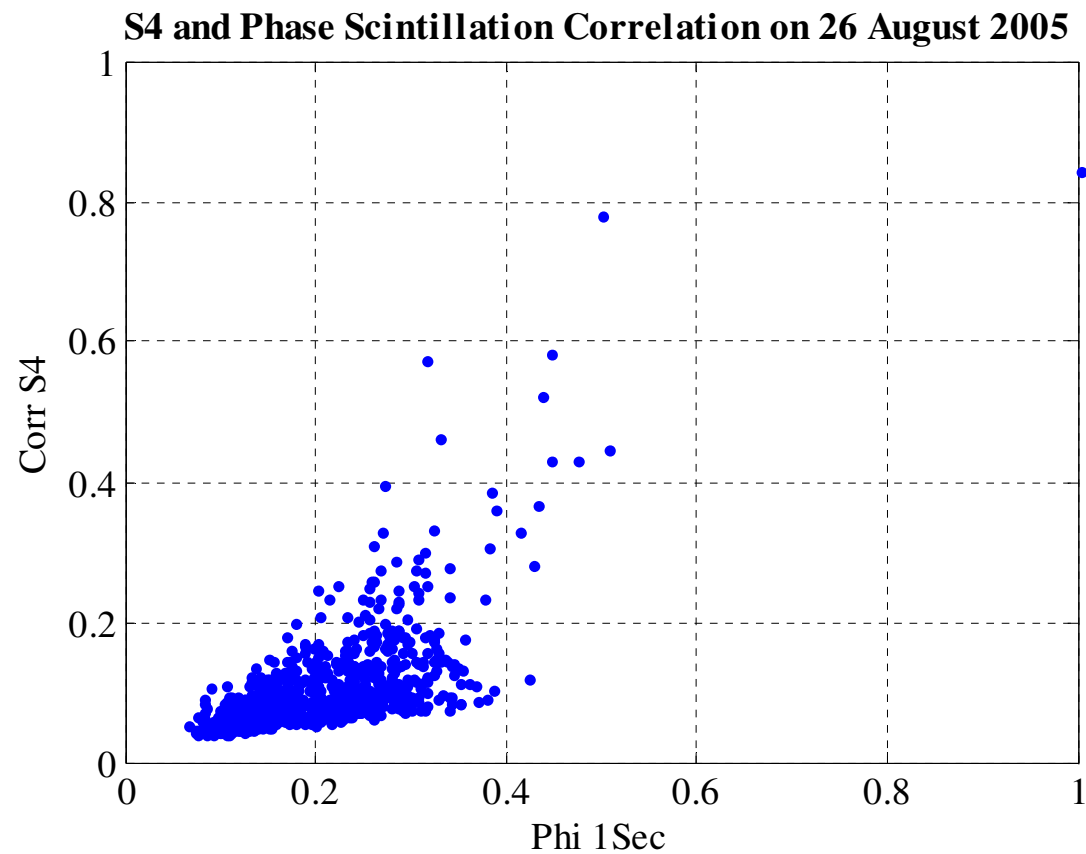
Moderate Scintillation



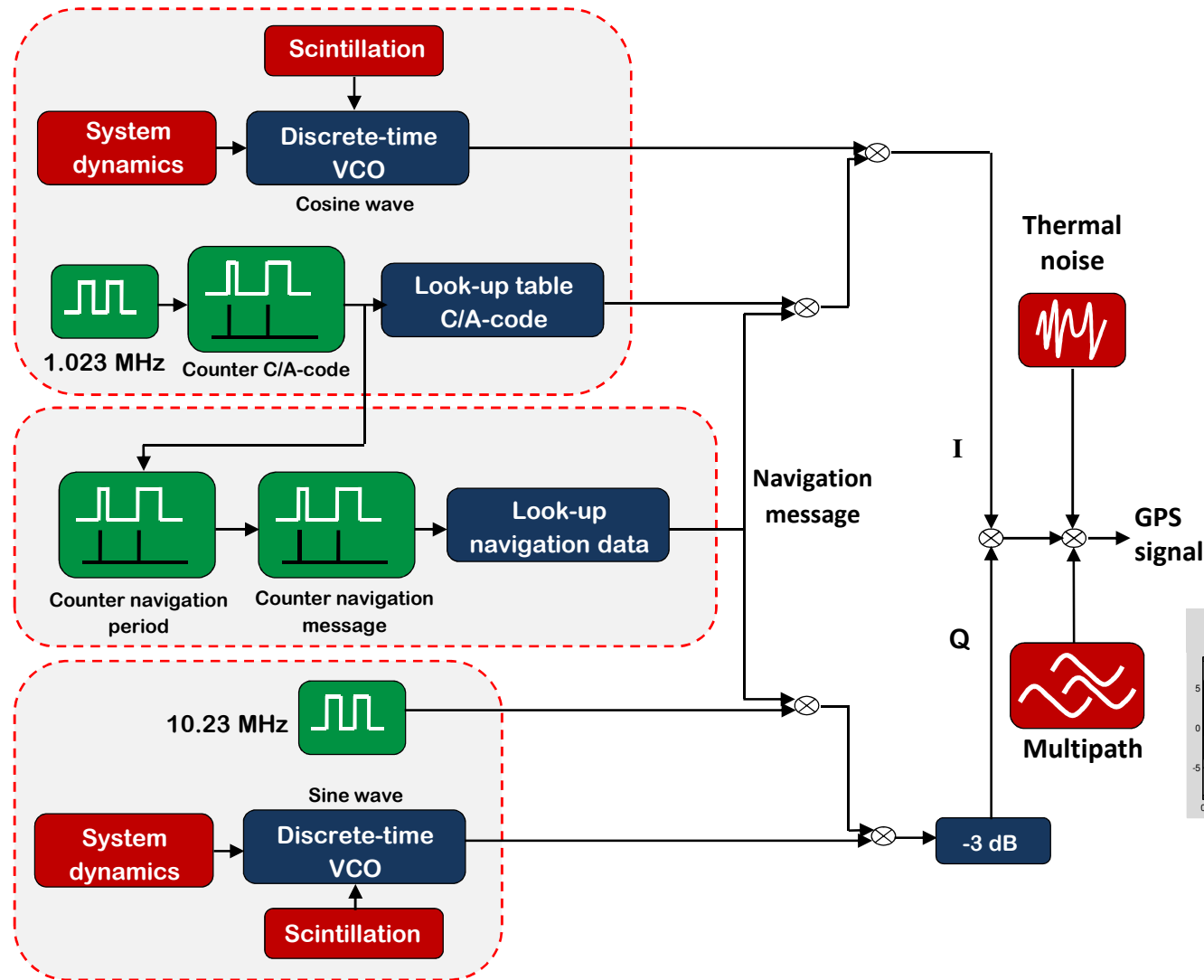
Strong Scintillation

Polar Scintillation Characteristics

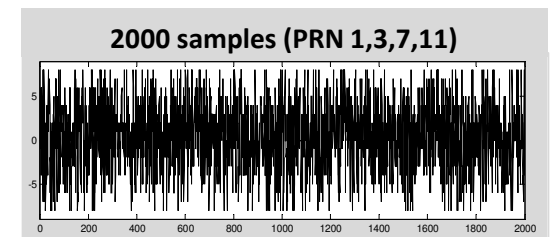
- Amplitude scintillation is significant
- S4 in range 0.4 to 0.8 for severe events
- Pairs (σ , S4) determined for scintillation simulations



GPS Signal Simulator



- Scintillation
- Thermal noise
- Multipath
- System dynamics
- **Conventional parameters for low-cost receivers**



Scintillation Simulator

Ionospheric **phase scintillation follows zero-mean Gaussian distribution** with standard deviation σ_ϕ .

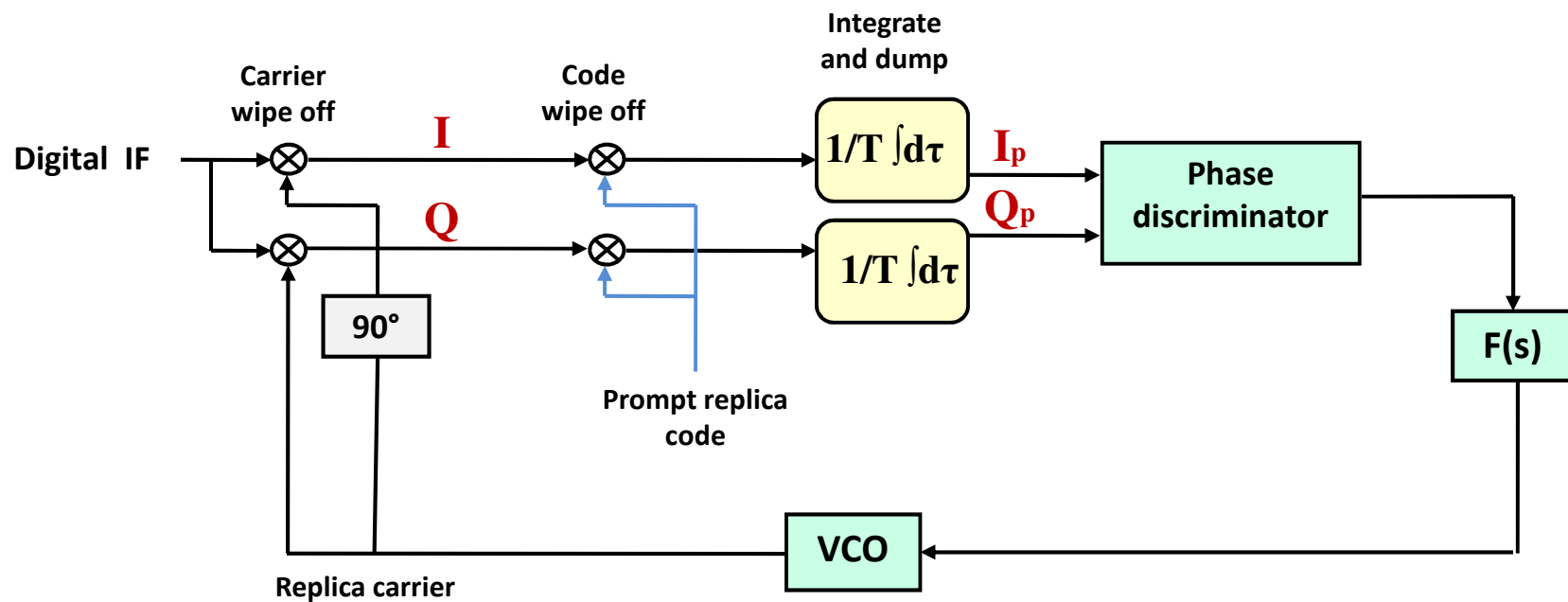
$$f(\phi) = \frac{1}{\sqrt{2\pi}\sigma_\phi} e^{-\phi^2/2\sigma_\phi^2}$$

Ionospheric **intensity scintillation follows Nakagami-m distribution** with normalized standard deviation S_4 . The intensity scintillation signal can be generated using **bivariate gamma function** $G(\alpha, \beta)$ with marginal density function given as follows:

$$f(\delta I = x) = \underbrace{\frac{m^m \delta I^{m-1}}{\Gamma(m) \Omega^m} e^{-m\delta I/\Omega}}_{\text{Nakagami-m}} = \underbrace{\frac{(x/\beta)^{\alpha-1}}{\beta \Gamma(\alpha)} e^{-x/\beta}}_{\text{Bivariate Gamma}} \quad \alpha = m = \frac{1}{S_4^2}, \beta = \frac{1}{m} = S_4^2, \Omega = 1$$

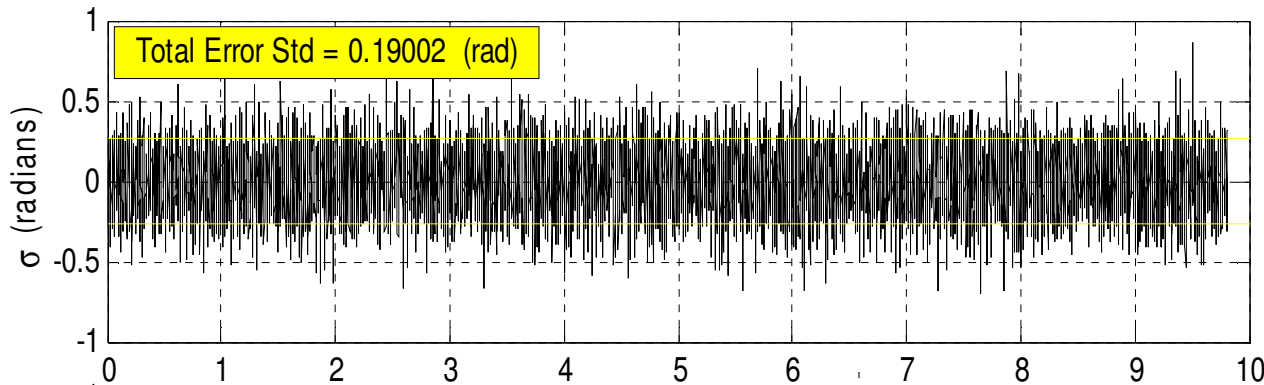
Tracking Loops

- Tracking loop bandwidth
- Pre-detection integration time



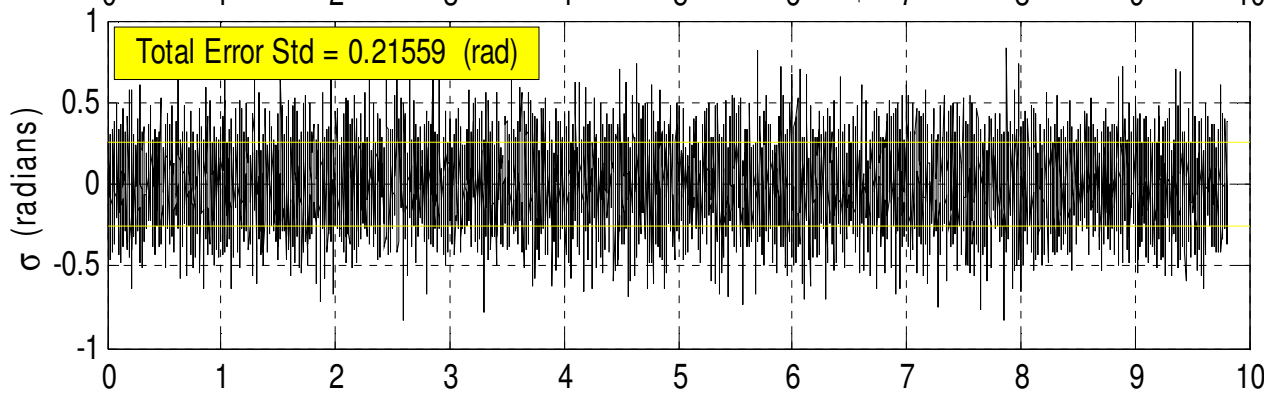
Generic Costas Carrier Tracking Loop

Carrier Tracking Error Auroral Region



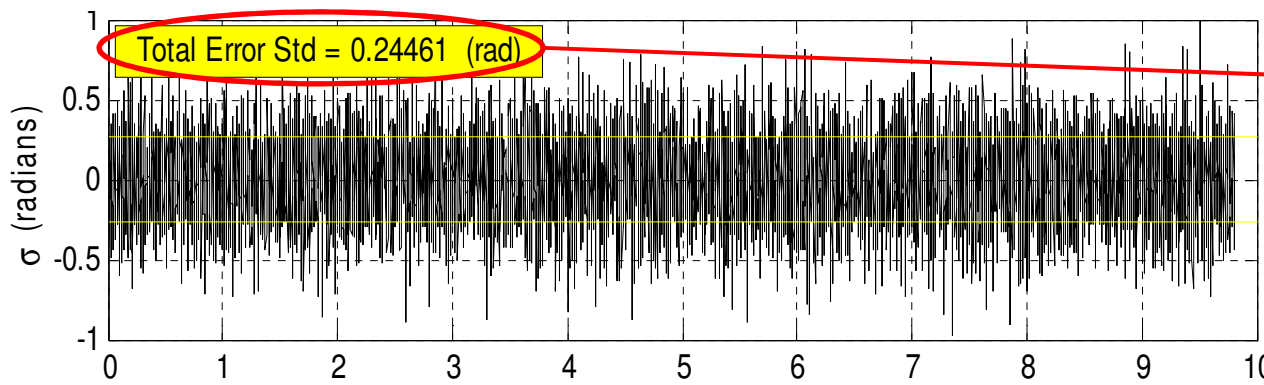
Weak Scintillation

$\sigma\phi_{60} = 0.1$ rad
 $S4 = 0.05$



Moderate Scintillation

$\sigma\phi_{60} = 0.5$ rad
 $S4 = 0.05$



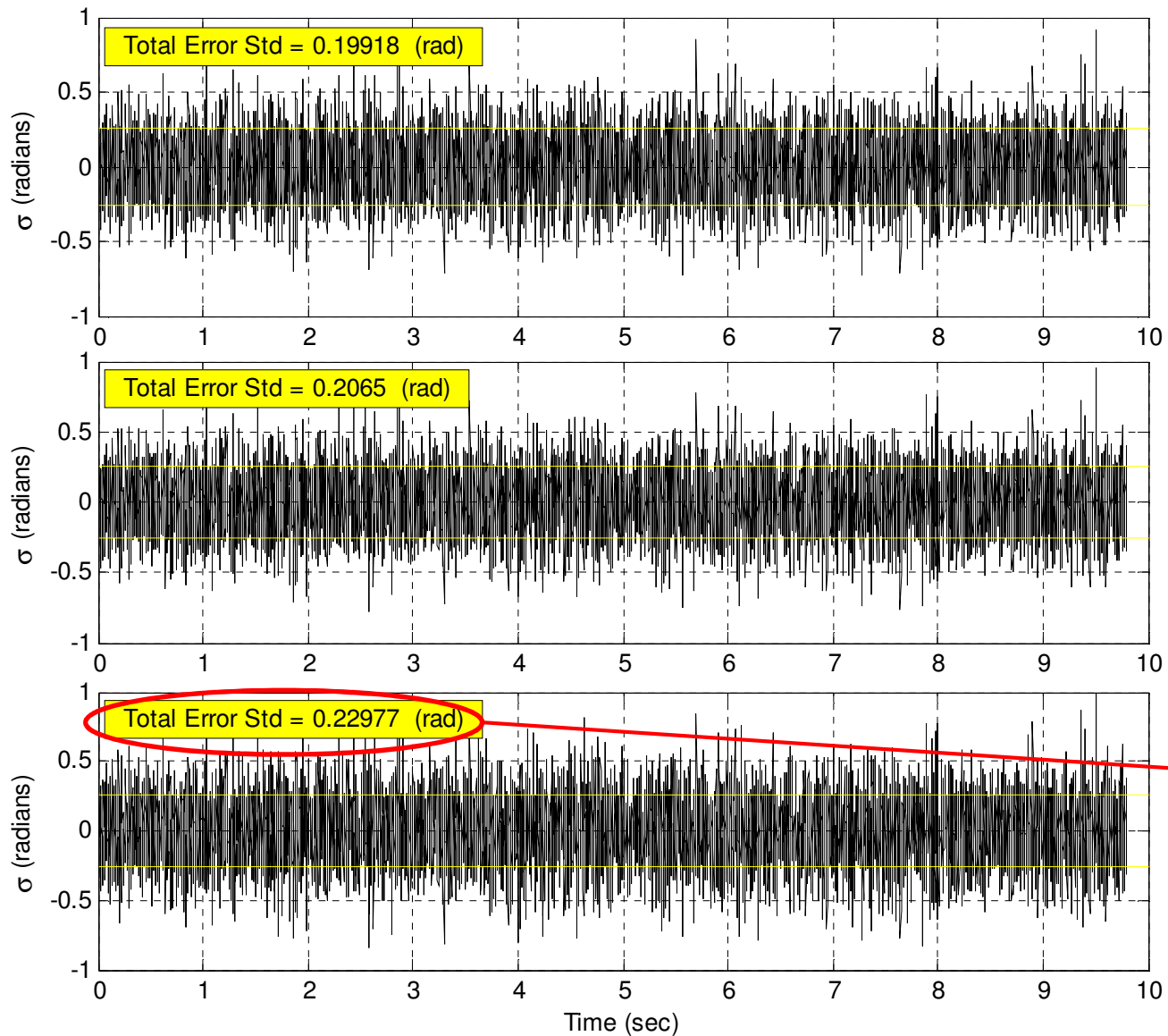
Strong Scintillation

$\sigma\phi_{60} = 0.7$ rad
 $S4 = 0.05$

Near the tracking threshold

$(B_n = 15$ Hz, $C/N_0 = 46$ dB-Hz)

Carrier Tracking Error Polar Region



Weak Scintillation

$\sigma_\phi = 0.2$ rad
S4 = 0.6

Moderate Scintillation

$\sigma_\phi = 0.4$ rad
S4 = 0.6

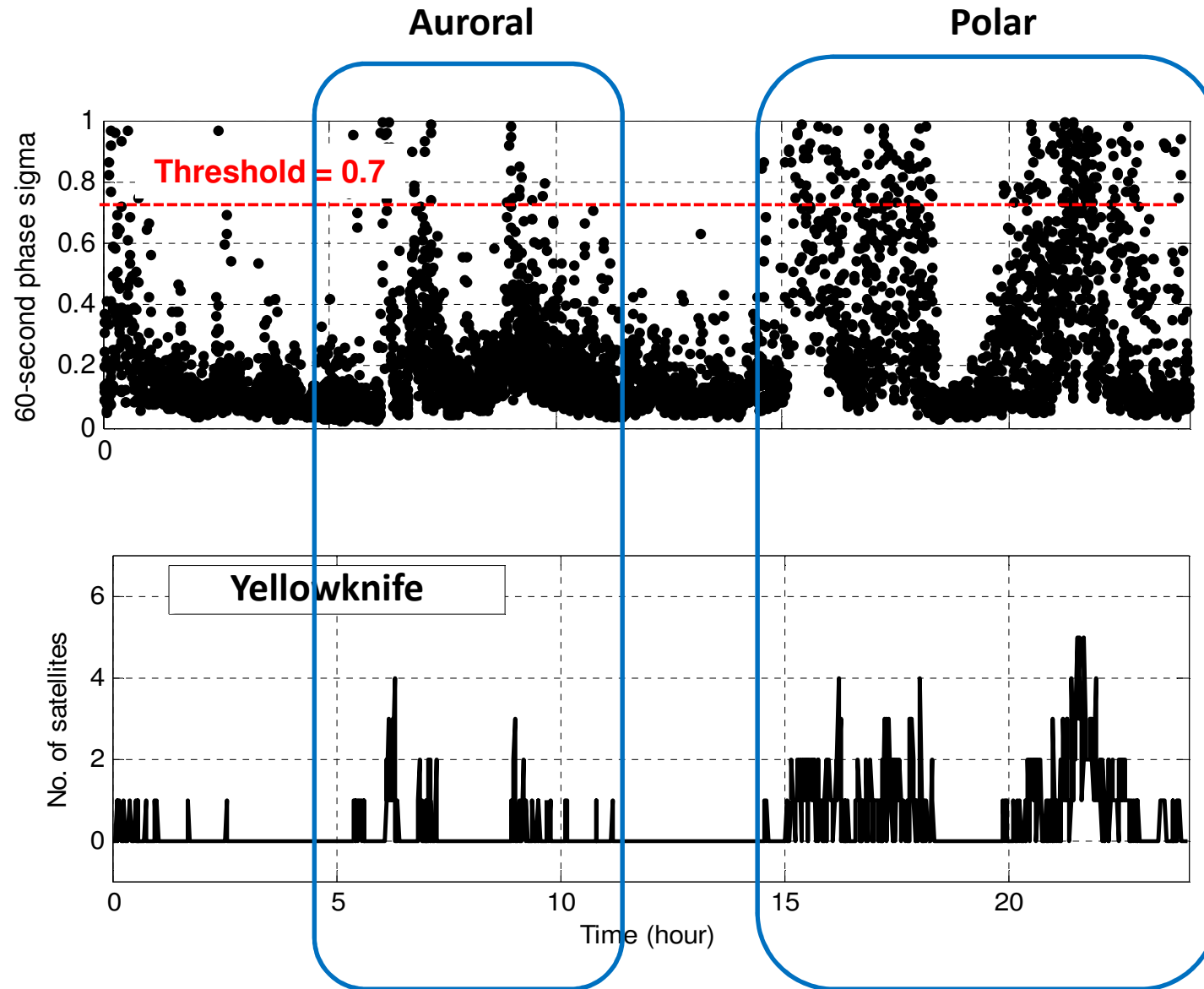
Near the tracking threshold

Strong Scintillation

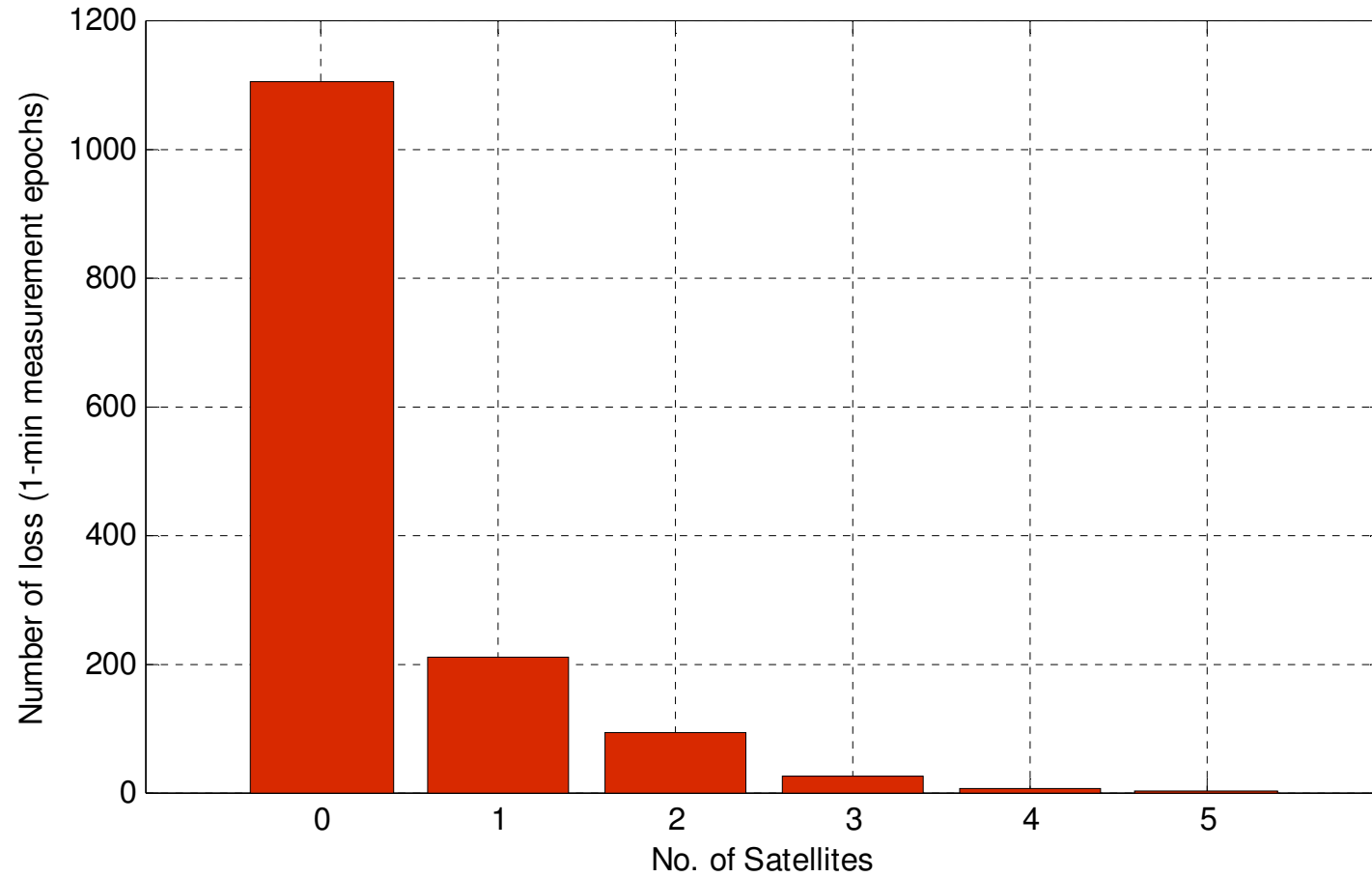
$\sigma_\phi = 0.6$ rad
S4 = 0.6

($B_n = 15$ Hz, $C/N_0 = 46$ dB-Hz)

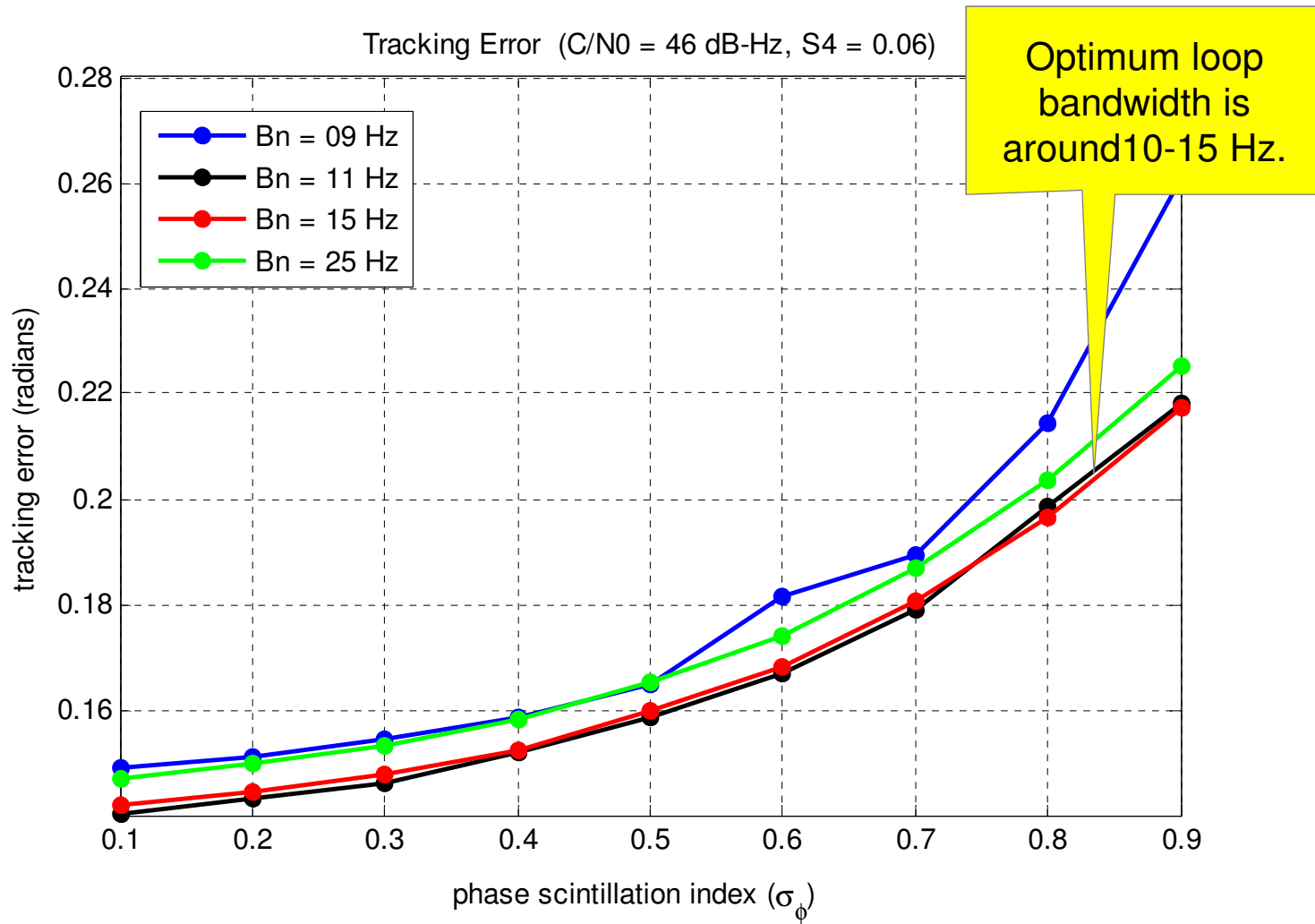
Example: Loss of Lock



Simultaneous Loss of Lock



Optimal Tracking Loop Bandwidth



($C/N_0 = 46$ dB-Hz, $S_4 = 0.06$)

Summary

- High latitude scintillations are due to different physical drivers (auroral versus polar regions)
- Simulations developed for auroral and polar regions
- Carrier tracking loop performance evaluated for generic low-cost receiver
- Receiver tracking loops may be challenged during severe scintillation